

# Derivation of Top-Of-the-Atmosphere Radiative Fluxes from SEVIRI: Methodology, Accuracy and Perspectives

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## 1. Introduction

Within the GERB data processing at RMIB, the solar and thermal fluxes at the TOA are, in a first step, estimated from the SEVIRI instrument. In a second step, those fluxes are “corrected” using the GERB instrument data.

The SEVIRI estimates are less accurate than the final fluxes. Nevertheless, these “un-corrected” fluxes may be of interest due to the difference of nominal life-span for the 2 instruments (SEVIRI=7 years > GERB =3 years).

This poster presents the methodology (measurement, NB-to-BB and radiance-to-flux conversions) to derive the SEVIRI fluxes estimates and their expected accuracy.

## 2. SEVIRI Measurements

- Temporal sampling 15' -> good resolution of the diurnal cycle, easy temporal integration.
- Field of view : 40% of the Earth (reduced to 20% if  $\theta_v$  is limited to  $60^\circ$ ).
- Use of Level 1.5 data (after EUMETSAT rectification) -> simplifies the processing (geolocation, clear sky quantities, ...).
- Noise level : depending on the channel but always small.
- Calibration error (this propagates up to the final product!):

Solar channels	(VIS0.6, VIS0.8, NIR)	5%	(vicarious calibration)
Thermal channels	(IR3.9, WV6.5, ..., CO2)	0.5%	(blackbody)

### 3. NB-to-BB conversion

In this step, the broadband unfiltered solar and thermal radiances must be estimated from the SEVIRI NB radiances:

$$\{ L_{nb} = \int (L_{sol}(\lambda) + L_{th}(\lambda)) \phi_{nb}(\lambda) d\lambda \} \quad - > \quad \begin{cases} L_{sol} = \int L_{sol}(\lambda) d\lambda \\ L_{th} = \int L_{th}(\lambda) d\lambda \end{cases}$$

This is done using second order polynomial regression on the SEVIRI NB radiances. Parameterization is done using best fit on a large database of realistic spectral signature  $L_{sol}(\lambda)$  and  $L_{th}(\lambda)$  generated using a radiative transfer model (SBDART).

- Good accuracy for thermal radiation (large number of channels).
- Limited accuracy for solar radiation (few channels and great dispersion of spectral signature).

## 4. Radiance-to-flux conversion - Solar

- Method : use of angular dependency models  $R(\theta, \phi)$  for the radiation at the TOA:

$$F_{sol} = \frac{\pi}{R(\theta, \phi)} L_{sol}(\theta, \phi)$$

- State-of-the-art models: CERES-TRMM ADMs (for about 200 scene types). About 30% more accurate than the ERBE models (depending on the kind of scene).
- ADM selection needs scene characterization in term of: surface IGBP geotype and snow cover, cloud fraction, cloud optical depth and cloud phase. In addition: the surface wind for the ocean sun-glint area.

## 5. Scene identification (only day time)

- Estimation of clear sky reflectance for the VIS0.6 and VIS0.8 channels (statistical estimation on the last 60 days),
- For each SEVIRI pixel, estimation of the cloud optical depth using the VIS0.6 (land) or VIS0.8 (ocean) channel and the corresponding clear sky value, thresholding on the cloud optical depth -> cloud mask at full resolution.
- cloud fraction estimated on 3\*3 pixels boxes,
- cloud optical depth averaged on the cloudy pixels in the boxes
- cloud phase estimated using the NIR1.6 channel.

## 6. Radiance-to-flux conversion - Thermal

- Method: use of limb darkening model  $R(\theta)$

$$F_{th} = \frac{\pi}{R(\theta)} L_{th}(\theta, \phi)$$

- Limb darkening model  $R(\theta)$  dependent on spectral signature in the thermal channels (does not need explicit scene identification):

$$R(\theta) = R(\theta, L_{6.2\mu}, \dots, L_{13.4\mu})$$

- Bottleneck: important overestimation of the OLR in case of optically thin ( $\tau \sim 1.5$ ) and cold (cirrus) clouds in the “nadir region” ( $\theta_v < 40^\circ$ ).

## 7. Accuracy table

steps	solar	thermal
NB measurement	5%	0.5%
NB-to-BB conversion	3.2%	0.7%
radiance-to-flux conversion	6 <sup>(1),(2)</sup> %	2 <sup>(1)</sup> %
<b>total</b>	<b>14.2%</b>	<b>3.2%</b>

Notes:

- (1). strong angular dependency of the radiance-to-flux conversion and hence regional pattern for the fluxes accuracy. The best accuracies are obtained for viewing angle close to 50° (e.g. Southern Europe),
- (2). this is **instantaneous error** using the CERES-TRMM ADMs (the temporal averaging reduces the angular conversion error). Preliminary validation has shown better results (4% instead of 6%, except for clear ocean).

## 8. Perspectives

### **Synergetic use of GERB data:**

- reduction of the measurement error and the NB-to-BB conversion error, especially for the solar radiation. GERB does not improve the radiance-to-flux conversion.
- Validation/improvement of the NB-to-BB conversion using the GERB data.
- Possible use of GERB data to calibrate the SEVIRI channels.

### **Synergetic use of CERES data (on the EOS satellites):**

- Validation of the SEVIRI derived fluxes (a software has been implemented at RMIB and tested using CERES-TRMM and Meteosat-7 data).
- Possibility to extend the spatial coverage to Northern Europe (CM-SAF activities).

## 9. Conclusions

- The methodology to derive TOA radiative fluxes from weather satellite imagers is well-known and has been widely applied (Meteosat, GOES, AVHRR, ...). No particular problem in the application to SEVIRI.
- Strong interest of GERB, especially for the solar flux estimation.
- Interest for dissemination to the user community only if GERB is not available. Nevertheless, interest to archive the SEVIRI fluxes for long-term analysis of the ERB.
- The method is currently running at RMIB using Meteosat-7 data (full disk and european sector).

## 10. Additional information

- Section “Documentation” at the RMIB GERB web site:

<http://gerb.oma.be>

- Near real-time TOA fluxes from Meteosat-7 available on the ftp site:

<ftp://gerb.oma.be>

- For the radiance-to-flux conversion for solar radiation see the CERES ADM group web site :

<http://asd-www.larc.nasa.gov/Inversion/>

- For the radiance-to-flux conversion for thermal radiation using SEVIRI spectral signature see: “Estimation of TOA Broadband Flux Errors Using a Coupled Radiometer-Imager. Part II: Longwave Radiation.” (to be published, available at RMIB).

- Spectral radiance curves at the TOA (solar and thermal) available at:

<http://gerb.oma.be/nic/SpectralRadianceTOA/>