

Extending TOA radiation to 1978

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Extending TOA radiation back to 1978 using wide field-of-view data

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Aim of this work

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Fill the gap!

Long time series of the total radiation emitted by the Earth, whose coverage extends as far as possible spatially and temporally, for the purpose of climate studies

- ► Total radiation (sw+Lw) emitted by the Earth, $W m^{-2}$
- Averaged per month
- Extent in time and space:
 - Coverage from +81° to -81° latitude
 - Coverage from -180° to +180° longitude
 - Coverage from November 1978 to September 1999 (nearly 21 years)



The wide field-of-view radiometer

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Nonscanner on board ERBS

- 2 wide field-of-view radiometers:
 - one radiometer for sw:
 0.2 to 5 μm
 - one radiometer for TW: all wavelengths

2 medium field-of-view radiometers

Source: http://mynasadata.larc.nasa.gov/images/erbenonscanner.gif



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Earth radiation budget



Source: http://mvnasadata.larc.nasa.gov/docs/earth radiation budget 17.pdf



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Justification

► Why wFov?

- Long, uninterrupted time series (scanner usually broke down after a few years)
- ▶ Why total radiation and not SW & LW?
 - sw filter suffers from ageing
 - Little degradation of the total radiation measurements over time
- Disadvantages
 - Cloud forcing cannot be studied
 - Low spatial resolution
- Compared to former work?
 - Improve accuracy of monthly averages by using state-of-the-art models and processing techniques



Data sources Experiments

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- ► ERB experiment on board NIMBUS-7 spacecraft
 - Solar and Earth Flux Data Tape (SEFDT) dataset
 - ▶ 1978/11 1993/01
- ERBE experiment on board ERBS, NOAA-9 and NOAA-10 spacecraft
 - Monthly Medium-Wide Data Tape (MWDT) dataset
 - ▶ 1984/11 1999/9





Data sources Spacecraft

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- ▶ NIMBUS-7
 - Sun-synchronous satellite with an inclination $\approx 99^{\circ}$
 - ► Local time equator passing, ascending node ≈ 11:00 (initially)
- **ERBS**
 - Precessing satellite with an inclination $\approx 57^{\circ}$
- NOAA-9
 - Sun-synchronous satellite with an inclination $\approx 99^{\circ}$
 - ► Local time equator passing, ascending node ≈ 14:30 (initially)
- NOAA-IO
 - Sun-synchronous satellite with an inclination $\approx 99^{\circ}$
 - ► Local time equator passing, descending node ≈ 7:30 (initially)



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- We assume the spectral response of the total wave (TW) measurement is sufficiently flat
- No intercalibration of different satellites (except common reference altitude)
- For the moment: albedo independent of solar zenith angle



Obtaining the monthly average flux

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Starting from instantaneous measurements from the wFov radiometer:

- Conversion of datafiles from native format to NETCDF
- Processing the raw measurements
- Binning in 5°× 5° bins
- Regression of the diurnal model on the data
- Numerical integration of the monthly average diurnal model from 0 to 24 hours
- Checks on the quality of the regression and final output



Conversion of datafiles from native format

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- These data were originally stored and processed on mainframes with tape drives
- Decode according to tape specifications
 - ERBE: User's Guide on http://eosweb.larc.nasa.gov/ GUIDE/dataset_documents/erbe_s7.html
 - ERB ON NIMBUS-7: NASA Contractor Report 170616 (Ray, Tighe & Scherrer, 1984)

 Add useful orbit & instrument information: inclination, orbit type (direct or retrograde), field-of-view aperture



Processing the raw measurements Overview

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- Add required quantities when not available in the datafiles
 - Solar zenith angle
 - Local time at nadir
- Direct sunlight elimination (geometric)
- Quality flags
 - Negative or otherwise unphysical data
 - Instrument looking off-nadir
 - Spacecraft status flags
 - Detector blinded by direct sunlight



Processing the raw measurements Elimination of direct sunlight

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- Not using solar zenith angle: seems to throw away good measurements without taking into account solar eclipse by Earth disc
- But using geometric technique
 - Takes into account angle between spacecraft, Earth and sun

 Takes into account solar eclipse by Earth (ellipsoid shape with GRS80 parameters)



Regression of the diurnal model

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- ► The diurnal model can be specified:
 - In terms of local time
 - In terms of (cosine of) solar zenith angle
 - Assume sloped line (daytime) intercepts flat line (night-time) at solar zenith angle of zero
- We've chosen to regress the model specified in terms of cosine of solar zenith angle
- Two-parameter diurnal model:

$$F(t) = \begin{cases} p_0 + p_1 \cos z(t) & \text{if } \cos z > 0\\ p_0 & \text{otherwise} \end{cases}$$
(1)

• Modified diurnal model during polar winter:

$$F(t) = p_0 \tag{2}$$

(日)



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 Solar zenith angle at noon varies considerably, depending on season and latitude

Regression of the diurnal model Why a diurnal model in terms of solar zenith angle?

- Corollary: even measurements at fixed local time lead to a range of zenith angles, which is better for the regression
- Regression of a linear two-parameter model is the obvious approach when only heliosynchronous data with two measurements per day are available
- Can estimate baseline (night-time) flux without night-time measurements

Disadvantage:

 Limited range of the independent variable at high latitudes



Regression of the diurnal model Example (Sahara region)





Numerical integration of the diurnal model Example (Sahara region)



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Checks on the quality of the regression Values can be rejected for several reasons

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- 1. An error occured during fitting (e.g., too many iterations in Levenberg-Marquardt)
- 2. Regression as a whole is not significant
- **3.** Regression is useless according to the Box criterion (explains less than the error)
- **4**. Null hypothesis cannot be rejected for at least one of the parameters
- **5.** At least one parameter is nonphysical (e.g., negative night-time flux)
- **6.** Numerical integration cannot be performed for numerical reasons
- 7. Resulting average flux is nonphysical (i.e., negative)



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- Fluxes as measured by the instrument at satellite altitude, but reduced to common altitude by inverse-square law
- ▶ ... And then mapped at nadir in boxes of 5°×5°
- First a set of maps of 1979, obtained using only NIMBUS-7 data (heliosynchronous, two measurements per day)
- Then a set of maps of 1987, obtained using all satellites (NIMBUS-7, ERBS, NOAA-9, and NOAA-10)





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80°

60°

40°

20°

20

-60°

-80°



Total radiation, W m⁻² September 1979



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Total radiation, W m⁻² September 1987



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Quality of the regression

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- Statistics for the regression
- R²: multiple correlation coefficient, test for linear correlation
- F: null hypothesis for all parameters simultaneously, test for significance of regression



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Quality of the regression December 1987



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Quality of the regression December 1987



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$\underset{R^{2}\text{-value}}{\textbf{Ouality of the regression}}$

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- ► No *R*² at high latitude in northern hemisphere: polar winter, one-parameter model
- Band of low-quality regression at 40 degrees latitude south
 - Seems to be caused by NIMBUS-7 (and sometimes NOAA-10) measurements
 - Related to refracted light?
 - Related to viewing geometry?
- R^2 seems to be higher over land than over ocean

- More pronounced diurnal cycle over land
- Related to cloud cover?



Quality of the regression



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An example of a good regression $R^2 = 0.98$



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Partially sunlit at sunrise/sunset transition $R^2 = 0.96$



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Zenith angle dependence over ocean surface $R^2 = 0.91$



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A very poor regression $R^2 = 0.05$ with NIMBUS-7 data, $R^2 = 0.96$ without



Estimate intercept with daytime measurements only $R^2 = 0.99$



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Modified diurnal model during polar winter $s^2 = 37705$ with NIMBUS-7 data, $s^2 = 91$ without



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Conclusions

- ► Revived the old NIMBUS-7 and less older ERBE data
- wFOV measurements do contain usable spatial information
- Made TOA radiation maps over nearly 21 years (November 1978 – September 1999) and nearly the entire globe, sometimes with scarce data

Problems remain: stray light? diurnal models inappropriate?



Future work

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- ► Incorporate more recent measurements (GERB, CERES)
- Better filtering of the data
- Improve diurnal models
- Applications: e.g. volcanic eruptions (El Chichón 1982, Pinatubo 1991)

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